

# Completeness study of the “train” sample

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# Rationale

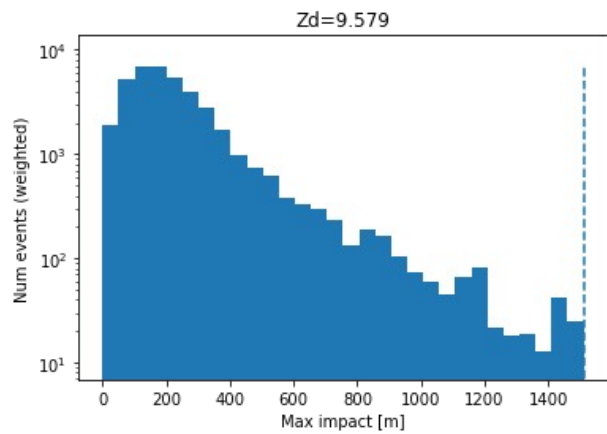
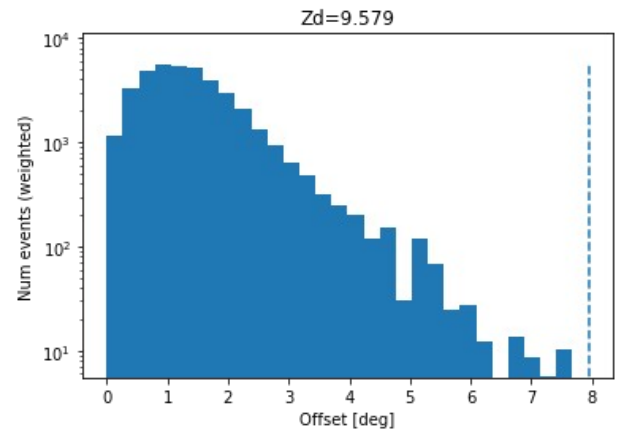
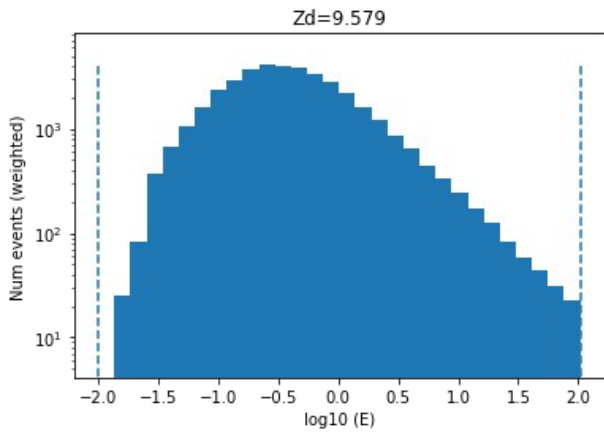
- The “train” samples were generated with a purpose of RF training – completeness (having nearly all the events that are possible to trigger the telescopes) was not a high priority
- We need protons (and Helium, electrons, ...) also for other purposes: MC sensitivity, matching trigger – there it is much more important to have the sample complete

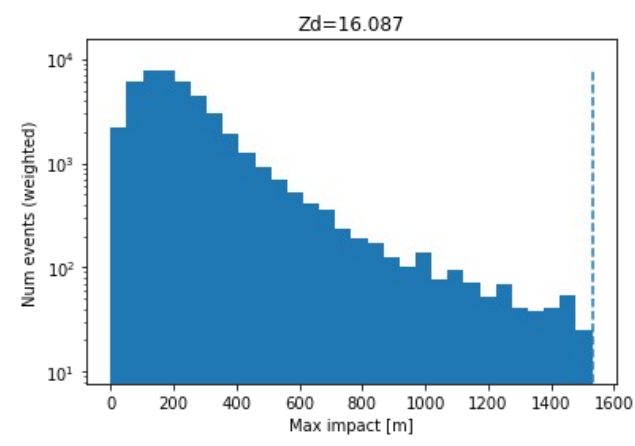
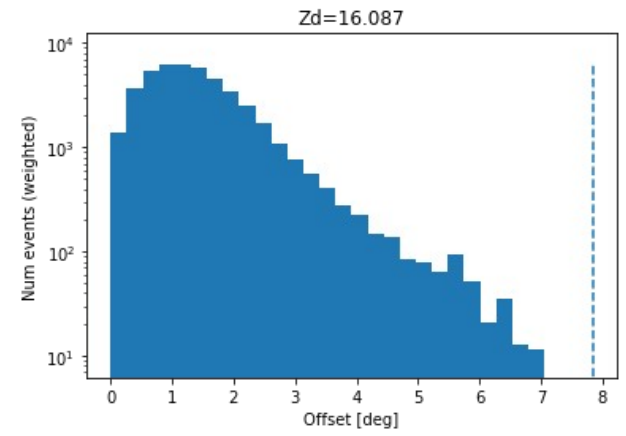
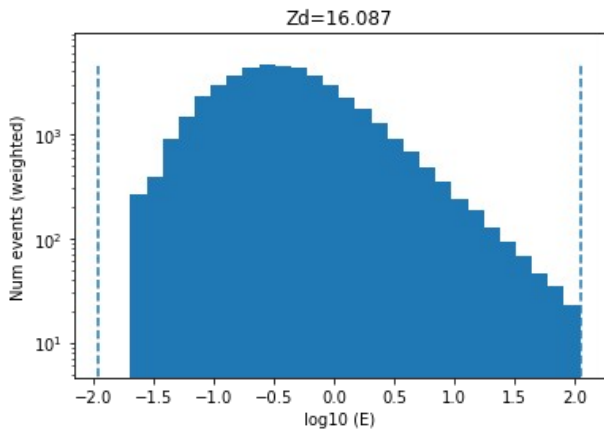
# Scalings of protons in the “train” sample

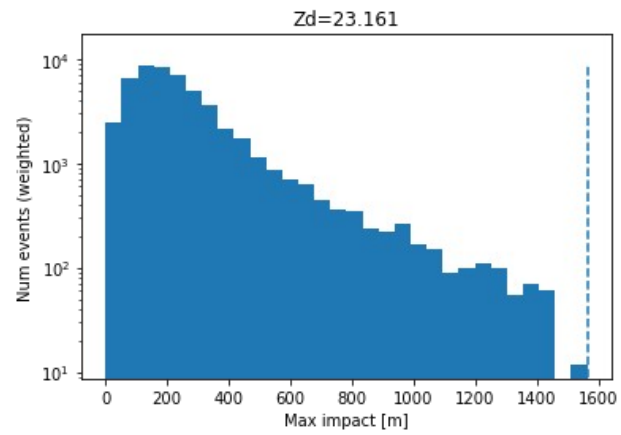
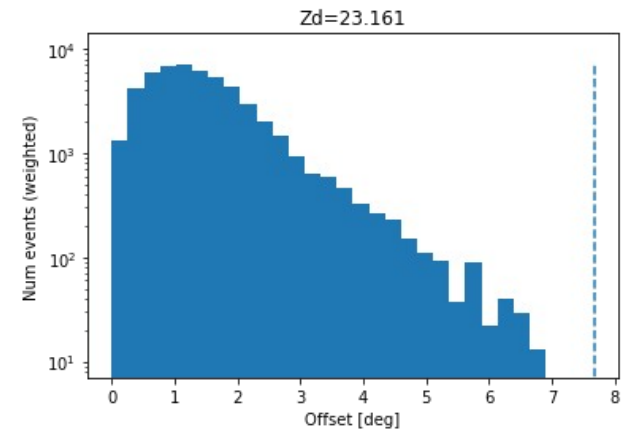
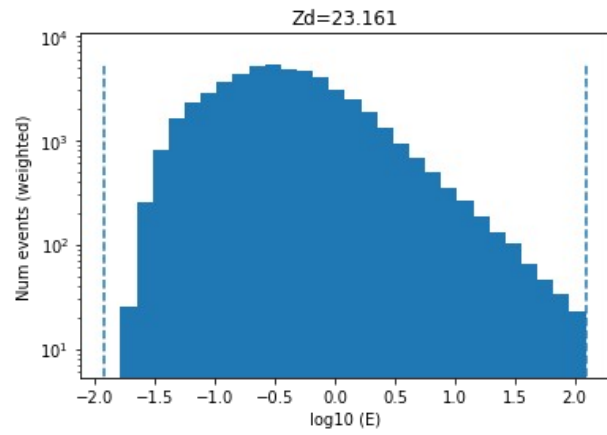
- <https://github.com/cta-observatory/lst-sim-config/issues/3>
- Viewcone: 8 deg scaled as  $\cos^{0.5} ZD$
- Energy range: 10GeV – 100 TeV scaled as  $\cos^{-2.5} ZD$
- Max impact: 1500 m, scaled as  $\cos^{-0.5} ZD$

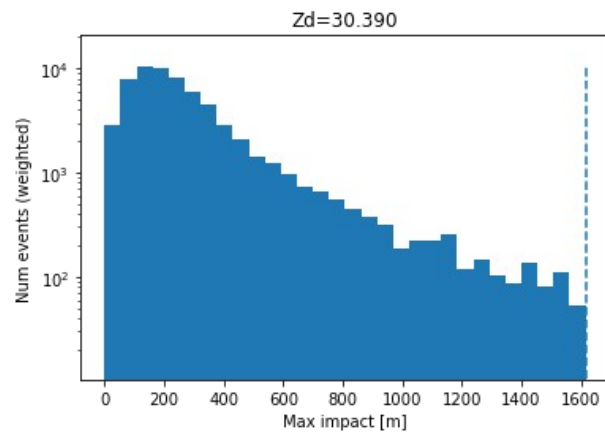
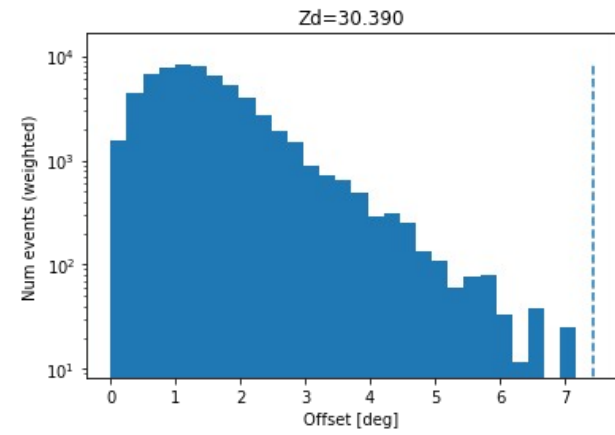
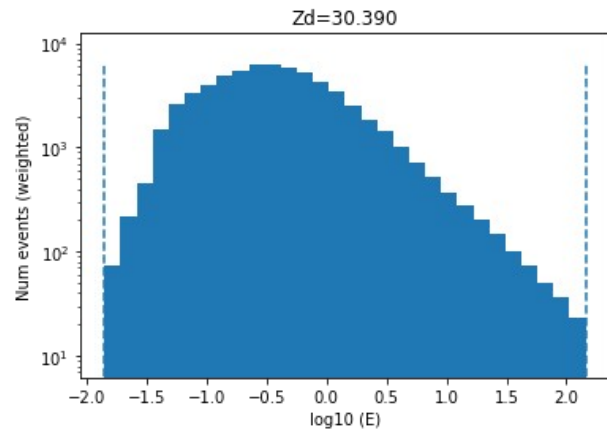
# Tests

- Using the train samples DL1 files I checked the basic completeness parameters (true energy range, impact, offset angle from the camera center).
- For impact there is actually a bit of asymmetry because LST1 is not at the center of the array.
- Plots are for LST1 telescope (the one with highest light yield – hence likely the most problematic one in terms of completeness), but with applied “magic\_stereo” condition (i.e. surviving the trigger condition of both LST1 and both MAGICs)
- Simulation slope is -2, but I weighted it to -2.8 to simulate the true proton spectrum
- Dashed lines show the simulated limits
- Each page is one zenith angle (stacking up a few azimuth angles)
- The range in all the panels is ( $y_{\max} / 1000$ ,  $y_{\max} * 2$ ), where  $y_{\max}$  is the maximum bin content (30 bins per simulated range)

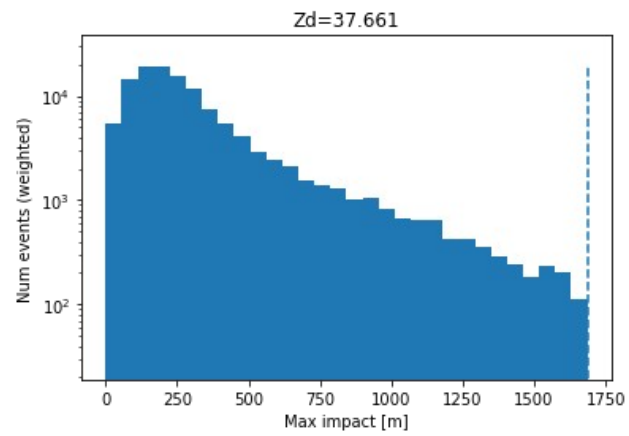
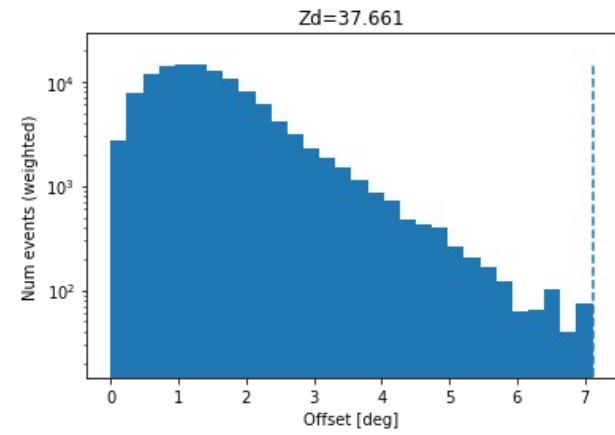
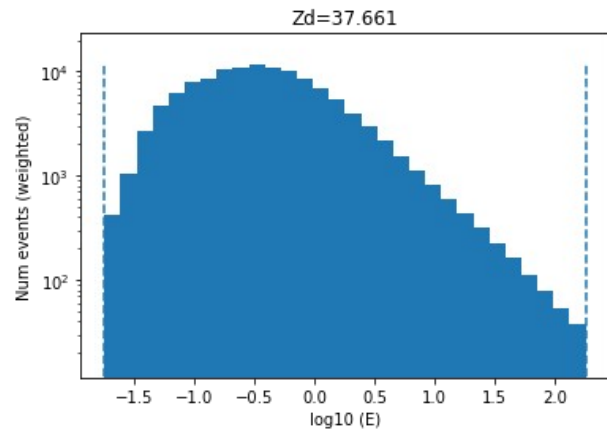


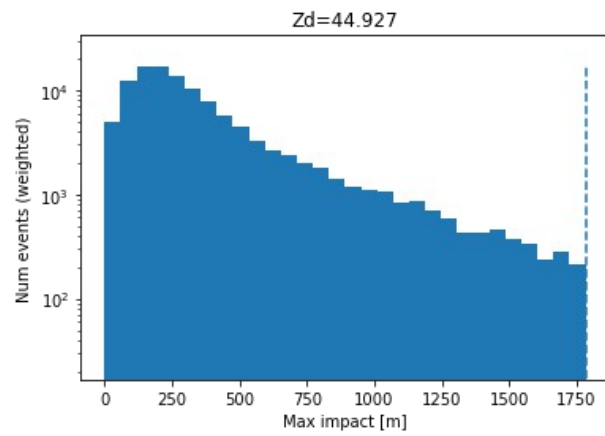
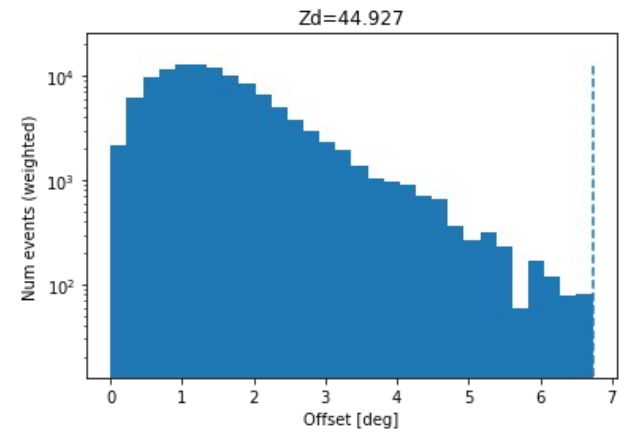
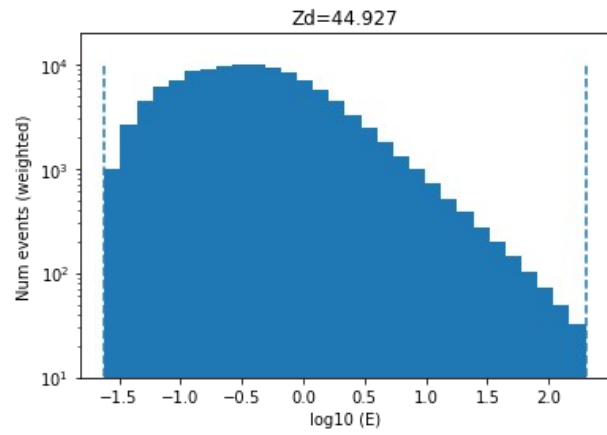


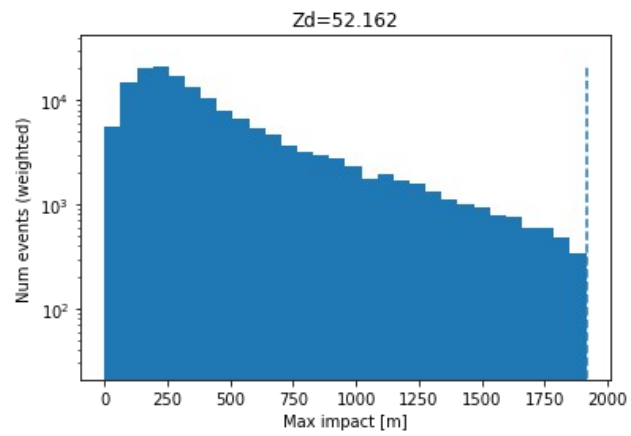
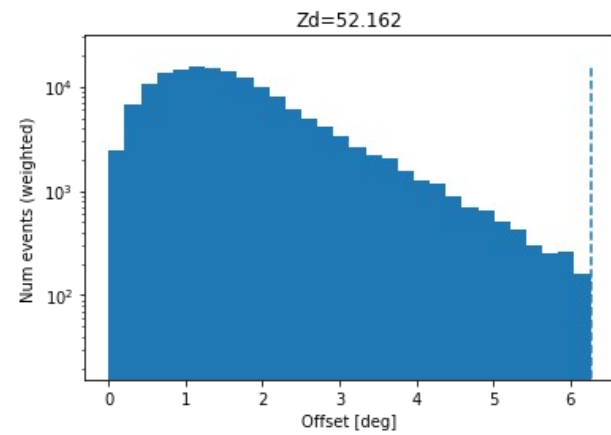
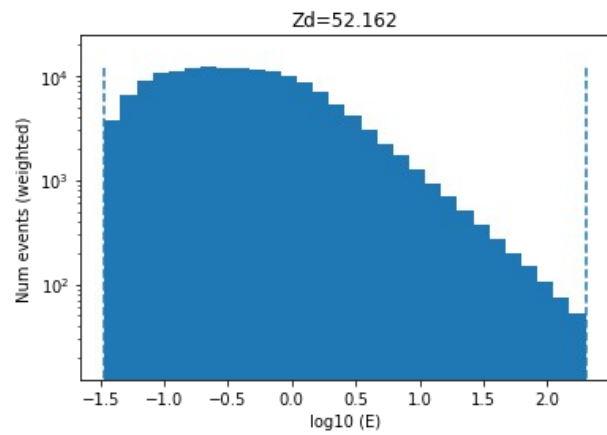


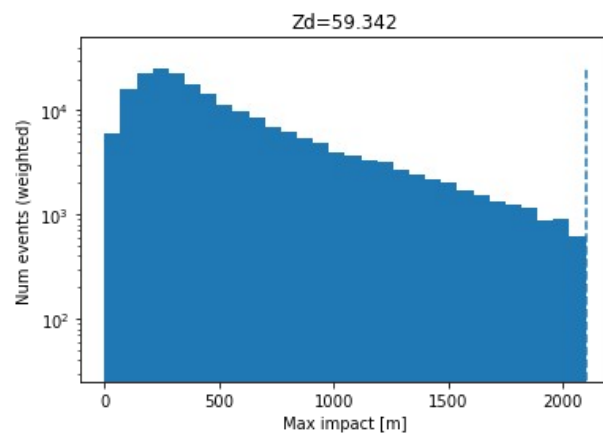
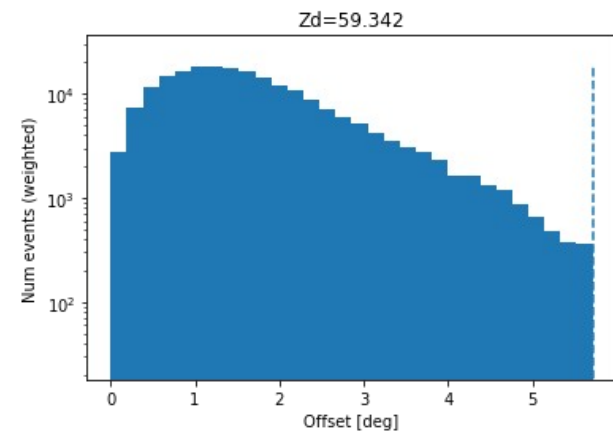
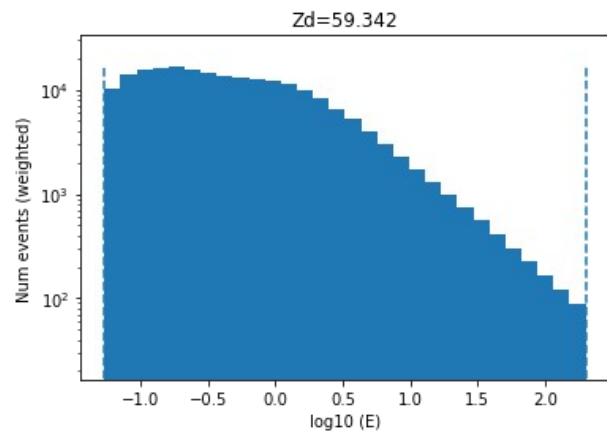


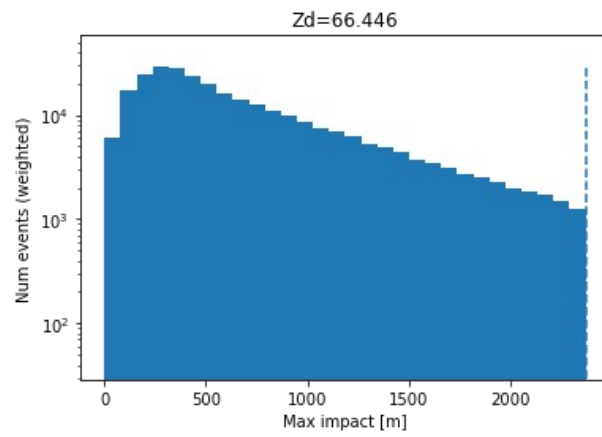
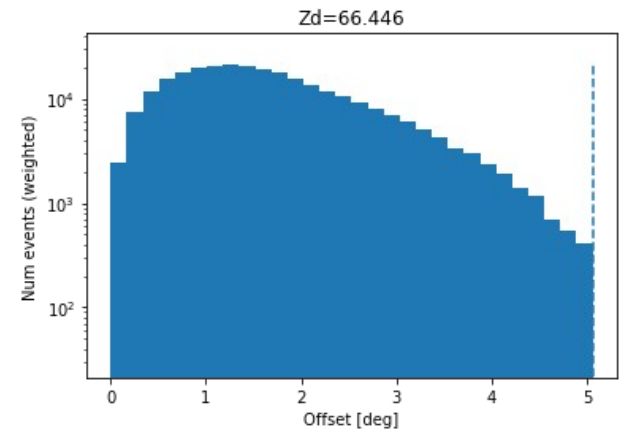
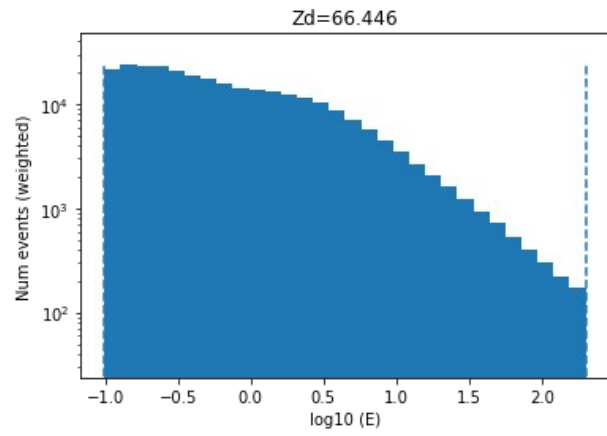












# Energy range

- While the  $\cos^{-2.5}ZD$  scaling might work for gamma rays, for protons it is much more complicated. The peak of the true energy distribution gets much broader with increasing zenith – effect of deep interacting events ? Single muons?
- Maybe apply a weaker scaling: e.g.  $\cos^{-1.5}ZD$ , this would result in 27GeV threshold instead of 52 GeV at the 59 deg zenith – highest energies are either way clipped at 200 TeV so will not be affected much

# Impact

- Scaling with  $\cos^{0.5}ZD$  definitely does not work, we are losing many high-impact events
- At 59deg we are simulating until 2 km and this makes a drop of an order of magnitude from the peak value at ~250m. Similar drop for 10deg zenith happens to distance of 800m
- I propose to use the “naive” scaling of  $\cos ZD$  instead (i.e. going to 3 km at 60 deg)

# Viewcone

- Shrinking the viewcone with raising zenith seems to be a bad idea, at low zenith 8 deg is perfectly fine (could be even slightly less), but the distribution is not getting narrower with increasing zenith but even slightly broader
- At low zenith the drop by an order of magnitude happens at 3.5 deg, at ZD=60deg it happens at 4.5 deg
- Propose to keep fixed 8 deg



# Summary

- At low zenith the samples look rather complete, but with increasing zenith it gets worse and worse.
- Proposed scalings of protons:
  - Energy range:  $(10\text{GeV}-100\text{TeV}) * \cos^{-1.5}ZD$
  - Impact:  $1500\text{m} * \cos ZD$
  - Viewcone: fixed 8 deg
- For helium:
  - Like protons, but multiply energy range by a factor of 2 (classical scaling is a factor of 4, but keeping a factor 2 of margin due to fluctuations at the lowest energies, the highest energies will be incomplete, but they are not that important for helium)